



Einstein proposed that three types of fundamental interaction were necessary to understand the interaction of light and matter under conditions of thermal equilibrium.

Absorption

















The linewidth of an atomic transition

- · Homogeneous broadening:
 - All atoms in the sample make the same contribution to the response at all frequencies;
 - Lorentzian lineshape function
 - e.g.
 - · Natural linewidth
 - · Collisional (pressure) broadening
- Inhomogeneous broadening
 - The atoms in the sample contribute differently to different frequency components of the total response;
 - Gaussian lineshape function
 - e.g.
 - · Doppler broadening
 - Broadening due to spatial variation in the environment (magnetic/electric field, the presence of localized perturbers)











$$g(v_0) = \frac{1}{\Delta v}$$

The absorption coefficient • To characterize the effect of finite atomic linewidth, introduce the lineshape function g(v) $\cdot g(v)$ is a function peaked around $v = v_0 = \frac{E_2 - E_1}{h}$ and $\int g(v) \cdot dv = 1$ Summing over all wavelengths: $\left(\frac{dN_2}{dt}\right)_{\text{stime}} = -\int g(v) B_{21} N_2 \rho(v) \cdot dv$ $\left(\frac{dN_1}{dt}\right)_{t=0} = -\int g(v) B_{12} N_1 \rho(v) \cdot dv$ With $g(v) = \delta(v - v_0)$ (Kronecker delta) we recover $\left(\frac{dN_2}{dt}\right) = -B_{21}N_2\rho(v_0)$

$$\left(\frac{dN_1}{dt}\right)_{abs} = -B_{12} N_1 \rho(v_0)$$

Let N = number per unit volume of photons N_v = number per unit volume of photons per unit frequency at frequency v, N_1 , N_2 = number per unit volume of atoms in ground and excited states $\frac{dN}{dt} = + \int g(v) \left(B_{21} N_2 - B_{12} N_1 \right) \rho(v) \cdot dv$ $=\int \frac{dN_{\nu}}{dt} \cdot d\nu$ where $\frac{dN_{\nu}}{dt} = +g(\nu) (B_{21} N_2 - B_{12} N_1) \rho(\nu)$ $= -g(\nu) \left(\frac{g_2}{g_1} N_1 - N_2 \right) B_{21} \rho(\nu)$ Using the spectral irradiance $I_{\nu} = \rho(\nu) c_n'$ and the identity $\rho(v) = N_v h v$ gives $\frac{dI_{\nu}}{dx} = \frac{n}{c}\frac{dI_{\nu}}{dt} = -g(\nu)\left(\frac{g_2}{g_1}N_1 - N_2\right)B_{21}\frac{nh\nu}{c}\cdot I_{\nu}$



Population inversion

We had for the small-signal absorption coefficient at frequency v:

$$\alpha(\nu) \equiv g(\nu) \left(\frac{g_2}{g_1} N_1 - N_2 \right) B_{21} \frac{nh\nu}{c}$$

If we can arrange for $N_2 - \frac{g_2}{g_1}N_1 > 0$

then there will be optical gain.

$$N_2 - \frac{g_2}{g_1} N_1$$
 is known as the population-inversion !

We can define a small-signal gain coefficient

$$\kappa(\nu) \equiv -\alpha(\nu) = g(\nu) \left(N_2 - \frac{g_2}{g_1} N_1 \right) B_{21} \frac{nh\nu}{c} \quad !$$

A laser exploits this gain to provide a highintensity beam of light.

Big assumption: the Einstein relations derived in conditions of thermal equilibrium also hold when there is population-inversion (negative effective temperature).

Example

Calculate the ratio of the ground state population to that of the excited state at room temperature, when the transition between the energy levels (which can be taken to be non-degenerate) corresponds to a transition in the visible spectral region.

Solution

Take the wavelength to be 550nm.

At thermal equilibrium

$$\frac{N_1}{N_2} = \frac{g_1}{g_2} \exp\left(\frac{h\nu}{k_B T}\right) = \frac{g_1}{g_2} \exp\left(\frac{hc}{k_B T \lambda}\right)$$

So
$$\frac{N_1}{N_2} = \exp\left(\frac{6.63 \times 10^{-34} \times 3.00 \times 10^8}{1.38 \times 10^{-23} \times 300 \times 550 \times 10^{-9}}\right)$$

$$\approx 9 \times 10^{37}$$





















Dipole emission patterns and scattering losses Individual atoms emit radiation with a dipole angular distribution Q. How can we reconcile this with the statement that for stimulated emission the emitted photon travels in the same direction as the incident photon? $\vec{k}_{emitted \ photon} = \vec{k}_{incident \ photon}$ A. Atoms throughout the gain medium emit stimulated photons that are coherent • The dipolar emission patterns of individual atoms interfere. • Provided the emitting atoms are uniformly distributed in space, the intereference of dipole emission patterns gives a resultant amplitude distribution that reproduces that of the incident beam.

• Spatial inhomogeneity of gain medium gives rise to scattering losses.

Population-inversion and the pumping threshold We had, for the gain coefficient: $\kappa(v_s) = \frac{N^* B_{21} nh v_s g(v_s)}{2}$ where $N^* \equiv \left(N_2 - \frac{g_2}{g_1}N_1\right)$ is the population inversion Writing $g(v_s) \approx \frac{1}{\Lambda v}$ and $B_{21} = \frac{c^3}{8\pi h v_s^3 n^3} A_{21} = \frac{c^3}{8\pi h v_s^3 n^3 \tau_{21}}$ gives $\kappa(v_s) = \frac{N^* c^2}{8\pi n^2 v_s^2 \tau_{21} \Delta v}$ so, at threshold $N^* = N_{th}^* = \frac{8\pi \, v_s^2 \, n^2 \, \tau_{21} \, \kappa_{th}(v_s) \, \Delta v}{c^2}$

Example Find the Doppler width of the carbondioxide laser transition at wavelength, λ =10.6 μ m, assuming the laser operates at 300K. Hence find the population inversion required to give a small-signal gain coefficient of 1 m⁻¹ for a carbon-dioxide laser, for which the Einstein Acoefficient of the upper laser level is 200 s⁻1 Find the pump power required to give the above value of the gain coefficient.

Solution Using $\Delta v_{Doppler} = \frac{2}{\lambda} \sqrt{\frac{2k_B T}{m}}$ $m_{CO_{\gamma}} = 44 \ a.u.$ gives $\Delta v_{Doppler} = 6.2 \times 10^7 Hz$ Then using $\kappa(v_s) = \frac{N^* B_{21} nh v_s g(v_s)}{c}$ with $B_{21} = A_{21} \frac{c^3}{8\pi h n^3 v^3}$ and $g(v_s) \approx \frac{1}{\Delta v_{Doppler}}$ $N^* = \frac{8\pi n^3 \kappa(v_s) \Delta v_{Doppler}}{A_{21} \lambda_s^2}$ Giving, finally $N^* = 7.0 \times 10^{10} \ cm^{-3}$



Gain saturation in homogeneously-broadened transitions

Consider the case in which one cavity mode lies in the spectral region of the active medium gain curve.

All atoms contribute to all parts of the medium gain profile

 \Rightarrow all atoms interact with the intra-cavity radiation

 \Rightarrow saturation decreases the entire gain profile





Rate-equation analysis

We now have expressions for the threshold gain and the population inversion at threshold.

The aim now is to find the condition that must be satisfied by the pumping mechanism to achieve the threshold population inversion.

To do this, we shall use a rate-equation analysis

This approximate treatment neglects:

• propagation effects on the phase of the intra-cavity light,

• quantum-mechanical superpositions of atomic energy levels



We have

$$\frac{dN_3}{dt} = R_3 - N_3 A_{32}$$
(1)

$$\frac{dN_2}{dt} = N_3 A_{32} - N_2 A_{21} - N_2 B_{21}\rho(\nu) + N_1 B_{12}\rho(\nu)$$

$$\frac{dN_1}{dt} = N_2 B_{21}\rho(\nu) - N_1 B_{12}\rho(\nu) + N_2 A_{21} - N_1 A_{10}$$

Solving for the steady-state and using the Einstein relation $g_2 B_{21} = g_1 B_{12}$, these equations transform to

 $N_{3} A_{32} = R_{3}$ $N_{2}[A_{21} + B_{21}\rho(\nu)] = N_{3} A_{32} + N_{1} \frac{g_{2}}{g_{1}} B_{21}\rho(\nu)$ $N_{2}[A_{21} + B_{21}\rho(\nu)] = N_{1} \left[A_{10} + \frac{g_{2}}{g_{1}} B_{21}\rho(\nu) \right]$ subtracting the last two yields $N_{1} = \frac{R_{3}}{A_{10}}$ Then $N_{2} = R_{3} \cdot \frac{1 + \frac{g_{2}}{g_{1}} \frac{B_{21}}{A_{10}}\rho(\nu)}{A_{21} + B_{21}\rho(\nu)}$

So the steady-state population inversion

$$N^* \equiv N_2 - \frac{g_2}{g_1} N_1 = R_3 \cdot \frac{1 + \frac{g_2}{g_1} \frac{B_{21}}{A_{10}} \rho(\nu)}{A_{21} + B_{21} \rho(\nu)} - \frac{g_2}{g_1} \frac{R_3}{A_{10}}$$

$$N^* = R_3 \cdot \left(\frac{1 - \frac{g_2}{g_1} \frac{A_{21}}{A_{10}}}{A_{21} + B_{21} \rho(\nu)} \right)$$

Below and exactly at threshold $\rho(\nu)$ is small.

If it can be neglected, we can write

$$N^* = \frac{R_3}{A_{21}} \cdot \left(1 - \frac{g_2}{g_1} \frac{A_{21}}{A_{10}}\right) \qquad \mathbf{!}$$

N* increases linearly with the pump rate below threshold

To achieve threshold we require a pumping rate $R_3^{th} = N_{th}^* \cdot \frac{A_{21}}{1 - \frac{g_2}{g_1} \frac{A_{21}}{A_{10}}}$ To achieve laser operation it is desirable that $\frac{A_{21}}{A_{10}} << 1$ which gives $R_3^{th} \approx N_{th}^* \cdot A_{21}$ Since $A_{21} = \frac{1}{\tau_{21}}$ we have, for the pumping power at threshold $P_{th} \approx N_{th}^* \cdot A_{21} \cdot (E_3 - E_0)$ We had $N_{th}^* = \frac{8\pi v_s^2 n^2 \tau_{21} \kappa_{th}(v_s) \Delta v}{c^2}$ The highest gain is produced at $V_s = V_0$. In this case $P_{th} = \frac{8\pi (E_3 - E_0) v_0^2 n^2 \kappa_{th}(v_0) \Delta v}{c^2}$









Plane-plane cavities are very lossy, so a laser cavity usually has at least one curved mirror. In such a situation, rays whose propagation is not restricted to the axi

Transverse modes

propagation is not restricted to the axial direction can still form stable, closed paths and there is a spectrum of transverse modes.

These modes can be properly described as Gaussian beams, which we shall meet later.

